SEARCHING FOR DARK PHOTON WITH THE PADME EXPERIMENT

C.TARUGGI ON BEHALF OF THE PADME COLLABORATION



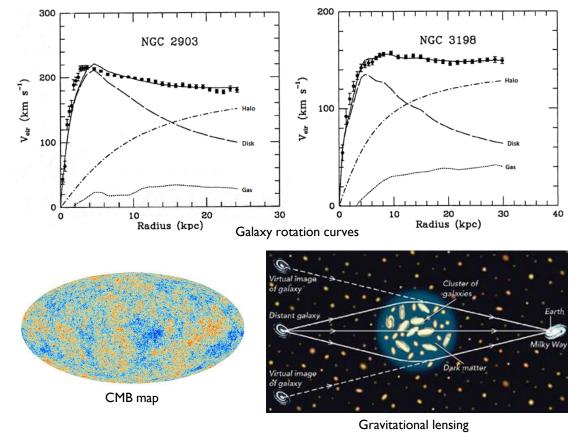
MATTER & GRAVITATION

What does the Universe tell us?

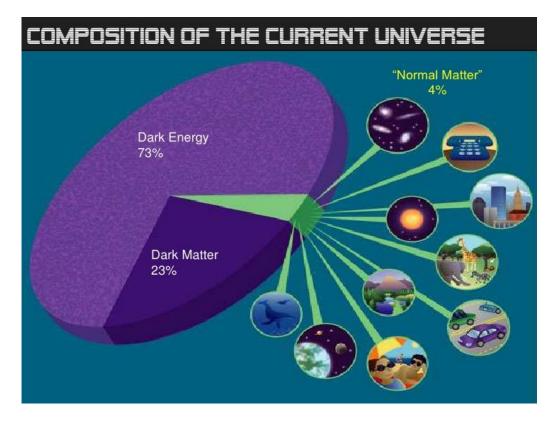
Visible matter cannot explain the behaviour of many cosmological phenomena (galaxy rotation curves, gravitational lensing, CMB distribution...). We can:

- I. modify gravitational laws
- 2. suppose the existence of another kind of matter, not producing radiation

Following the second hypothesis, how can we detect such an «invisible» matter? Does it interact with Standard Model (SM) particles, and how?



INTRODUCING DARK MATTER



If gravitation law is right, current observations tell us that we know only $\sim 4\%$ of the Universe.

We suppose that dark matter (DM) is about 23% of the total, and that interacts *at least* gravitationally with SM particles, but very weakly! DM could belong to the SM particles group (e.g. WIMP), but could also be something else.

What if DM and SM particles interact through another force?

This could explain why detection of DM is so difficult: DM could live in a different world with respect to SM particles, and these two worlds could be connected by a fifth interaction.

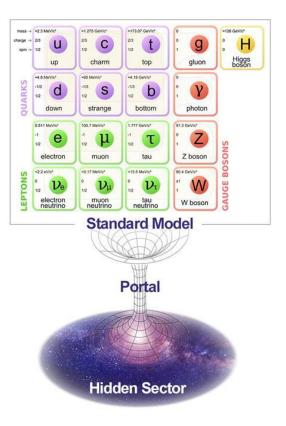
DARK SECTOR

DM could live in the so called *dark sector*. We can introduce a new interaction that connects ordinary matter with the dark sector: SM particles could be neutral under this interaction. A mediator particle is associated to the interaction. Depending on the model, the mediator can be a fermion, a boson, a vector, etc.



Dark sector theories could also be used to explain other physical phenomena other than DM, for example the muon g-2 anomaly and the ⁸Be anomaly.

The simplest dark sector model introduces a U(1) symmetry, whose mediator is a vector boson A': in analogy with the standard photon, we call it *dark photon (DP)*.



DARK PHOTON: PRODUCTION AND DECAY

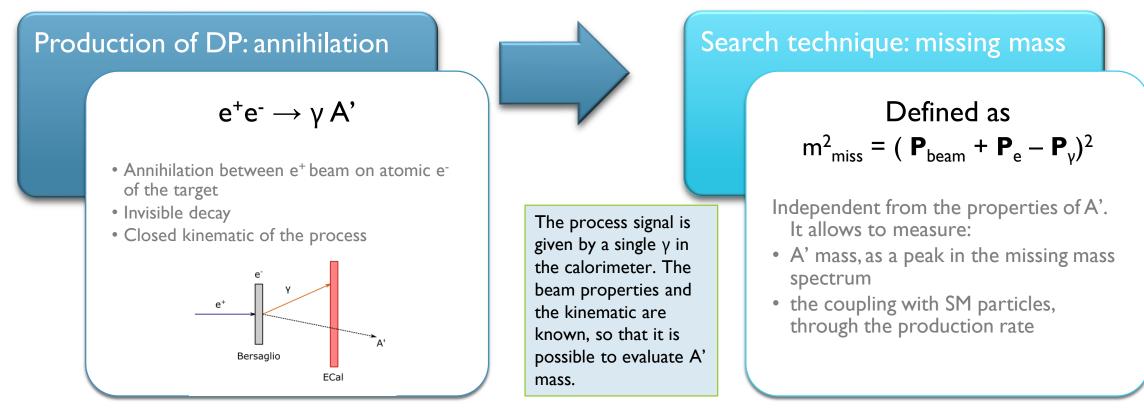
According to the dark sector model one considers, there are different ways dark photon can be produced and decay.

Here we consider production from e^+e^- interaction.

Production			Decays	
ANNIHILATION	BREMSSTRAHLUNG	MESON DECAY ^(*)	VISIBLE	INVISIBLE
e SrA.	_+ ► A'		(SM particles)	(DM PARTICLES)
e ⁺ V e ⁺ V	e v		If DM particles with m _{MO} ≤ m _{A'} /2 do NOT exist	If DM particles with m _{MO} ≤ m _{A'} /2 DO exist

Searching for visible or invisible decays lead to different dark photon properties (e.g. mass and coupling), meaning that different exclusion plots are obtained.

DETECTION OF DARK PHOTON IN PADME

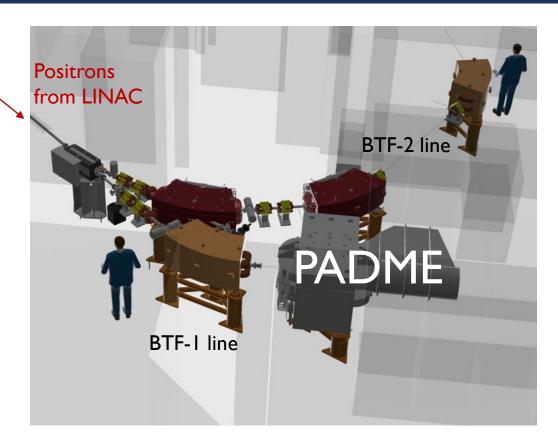


XIX LNF SPRING SCHOOL "BRUNO TOUSCHEK"

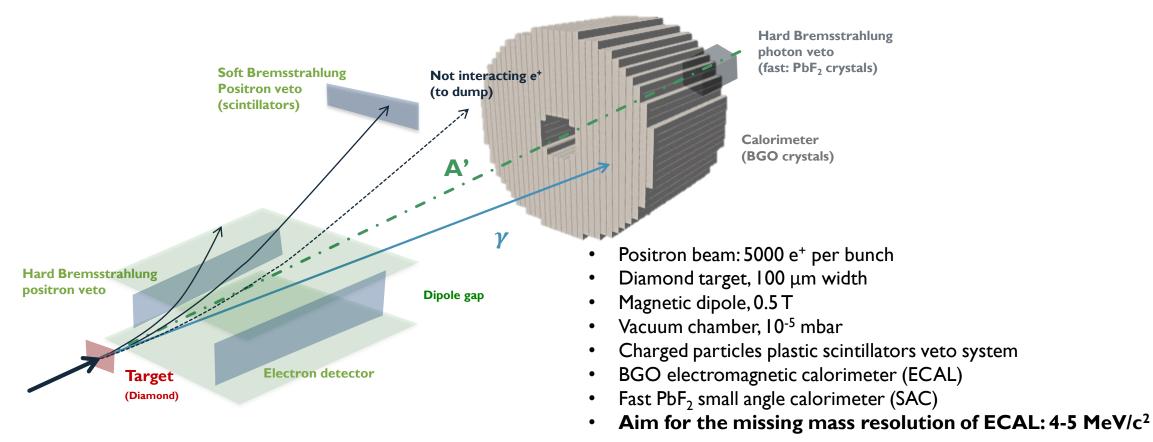
THE BEAM TEST FACILITY

	electrons	positrons		
		•		
Maximum beam energy (E _{beam})	750 MeV	550 MeV		
Linac energy spread	0.5%	۱%		
Typical Charge	2 nC	0.85 nC		
Bunch length	I.5 – 40 ns			
LINAC repetition rate	I-50 Hz	I-50 Hz		
Typical emittance	I mm mrad	~1.5 mm mrad		
Beam spot size	<i mm<="" td=""></i>			
Beam divergence	I-I.5 mrad			

PADME will be placed in the LNF Beam Test Facility (BTF).



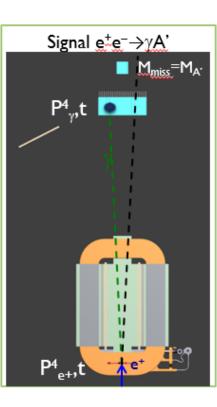
MAIN FEATURES OF PADME



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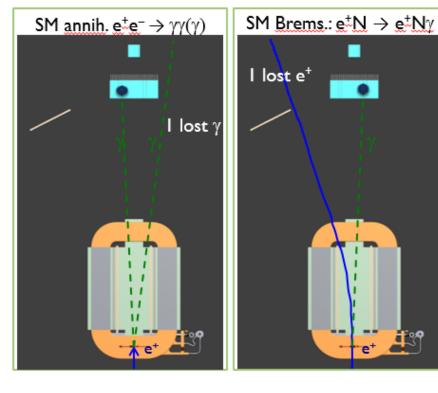
SIGNAL AND BACKGROUNDS

Signal: Ιγin ECAL



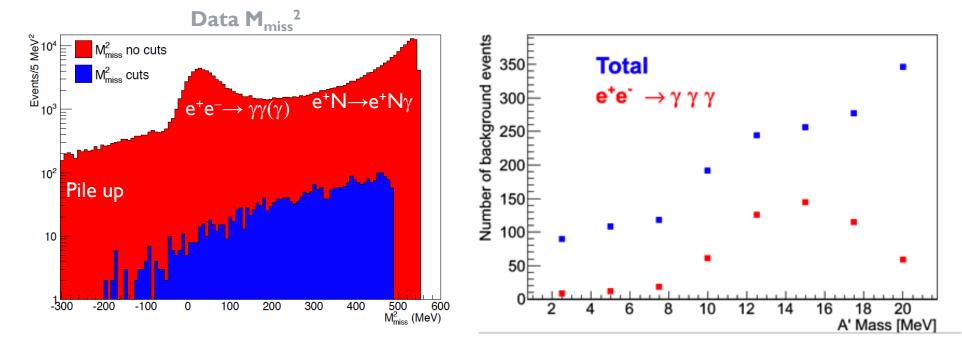
Backgrounds





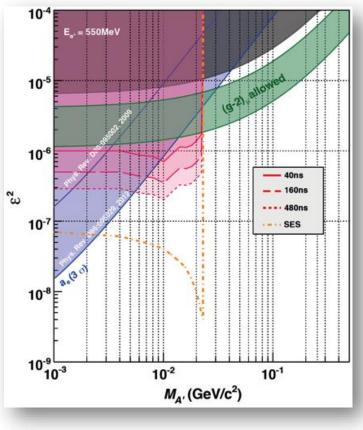
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BACKGROUND ESTIMATE



- Main background contribution given by: $e^+e^- \rightarrow \gamma\gamma$, $e^+e^- \rightarrow \gamma\gamma(\gamma)$, $e^+N \rightarrow e^+N \gamma$, pile up
- Background from pile up is important, but rejected by the maximum cluster energy cut and M_{Miss2}.

SENSITIVITY STUDIES



- Sensitivity studies have been performed using a GEANT4 simulation
- 2.5 · 10¹⁰ 550 MeV e⁺ on target events extrapolated to 10¹³
- PADME can explore in a *model-independent way* the band up to M²_A, = 2m_eE_{e+} (red band in the plot)
- $E_{e+}=550 \text{ MeV} \rightarrow M_{A'} < 23.7 \text{ MeV/c}^2$

CONCLUSIONS

- The PADME experiment will search for dark photon, as one of the possible mediator of a fifth interaction that connects our world to the so called dark sector, where DM lives
- PADME will study the invisible decay of DP, using the annihilation process $e^+e^- \rightarrow A'\gamma$
- The experiment will search for a DP with mass up to 23.7 MeV (beam energy = 550 MeV)
- The data taking will start in June 2018 at LNF
 - We expect to collect ~1x10¹³ positrons on target by the end of 2018
 - The main components of the experiment are ready
 - Expect a second physics run in 2019 at LNF